

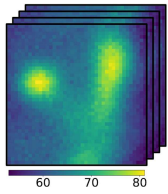
Multispectral and Hyperspectral Image Fusion with JWST/MIRI

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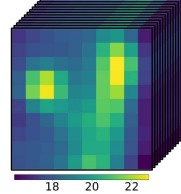
Context

• **Data** : multispectral cube acquired by an imager, hyperspectral cube by a spectrometer



Multispectral data* y_m

- High spatial resolution
- Low spectral resolution



Hyperspectral data* y_h

- Low spatial resolution
- High spectral resolution

• **Objective** : restoration of a cube x with **enhanced** spatial and spectral resolutions.

• **Hypotheses** :

• Subspace approximation with the Linear Mixing Model [3] [4], such that

$$x[i, j, l] = \sum_{t=1}^T a_t[i, j] s_t[l]$$

$$x = T a$$

• Data corrupted with additive white gaussian noises.

• The imager and spectrometer models are known [3] [4].

*simulated observations of the Orion Bar [1] [2]

Methodology

• Case of an **ill-posed inverse problem**, solved by minimizing a regularized convex criterion

$$\hat{a} = \underset{a}{\operatorname{argmin}} \left\{ \underbrace{\mu_m \|y_m - M a\|_2^2}_{\text{Data adequation}} + \underbrace{\mu_h \|y_h - H a\|_2^2}_{\text{Regularization}} + \underbrace{\mu_r R(a)}_{\text{Regularization}} \right\}$$

• Two regularizations used : quadratic (ℓ_2 -norm) half-quadratic ($\ell_{2,1}$ -norm) [6]
 $R(a) = \|D a\|_2^2$ $R(a) = \varphi(D a)$

• Both cases : resolution of a linear system $Q \hat{a} = q$, solved in the literature [5] with gradient based algorithms for the ℓ_2 -norm, where

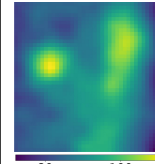
$$Q = \mu_m M^H M + \mu_h H^H H + \mu_r D^H D$$

Contribution

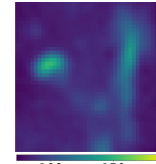
• Proposed procedure for the **fast** and **exact** calculation of Q^{-1} by demonstrating its diagonal block structure using [7] and applying a matrix inversion method from [3].

• Two main contributions :

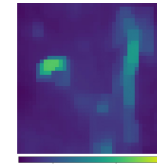
- the fast calculation of the **exact solution** for ℓ_2 , with $\hat{a} = Q^{-1} q$.
- an **accelerated** procedure for the alternating minimization problem [3][4] for $\ell_{2,1}$.



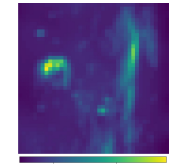
Coaddition



Exact solution of ℓ_2 [5]



Proposed $\ell_{2,1}$ approach



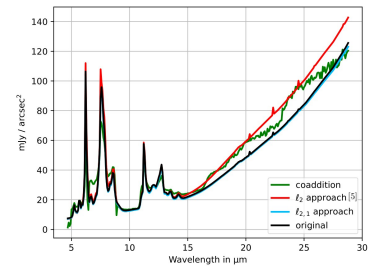
Original

Methods	NRMSE ($\times 10^{-3}$)	dSSIM ($\times 10^{-5}$)	SAM ($\times 10^{-3}$)	PSNR	Time [s]
Coaddition	133	1476	119	37	0.6
Exact solution of ℓ_2 [5]	27	241	5.8	50	2 (with prep.)
Proposed $\ell_{2,1}$ approach	22	179	4.0	52	19 (300 iter.)

• Efficient **deconvolution** and **denoising** for all wavelength with inverse problem approaches, mainly thanks to correlations induced by the Linear Mixing Model

• Exact solution of ℓ_2 **1000 times faster*** than minimization with gradient based algorithm [5] for a low noise case (SNR = 100 dB)

• Best spatial and spectral resolutions found with the proposed edge-preserving $\ell_{2,1}$ approach



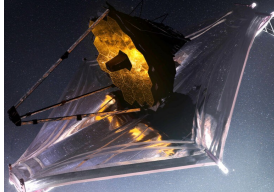
*Size MS dataset : 9 x 124 x 248, Size of HS dataset : 300 x 31 x 62, Size of reconstruction : 300 x 124 x 248.

Instrument models

• Application to the Mid-InfraRed Instrument (MIRI) of the James Webb Space Telescope (JWST)

• MIRI has 2 components :

- an imager, MIRIM,
- an integral field spectrometer, the MRS.



JWST

• Spectral range of MIRI : 5 to 28 μm .

• **Imager model**

• Spectrally varying spatial blur, i.e. convolution with the imager impulse response h_m

• Spectral response of the imager w_m

• Spectral integration over C bands

• Forward model $\Rightarrow y_m^c[i, j] = \sum_l (x * h_m)[i, j, l] w_m^c[l] + n_m^c[i, j]$

$$y_m = \underbrace{W_m C_m T a}_{= M, \text{ forward imager model matrix}} + n_m$$

• **Spectrometer model**

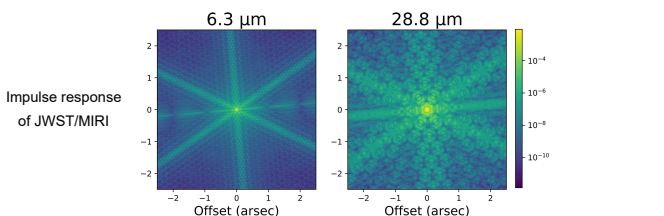
• Spectrally varying spatial blur, i.e. convolution with the imager impulse response h_h

• Spectral response of the spectrometer w_h

• Spatial subsampling (aliasing)

• Forward model $\Rightarrow y_h[\bar{i}, \bar{j}, l] = \sum_{i=\bar{i}d_i}^{(\bar{i}+1)d_i} \sum_{j=\bar{j}d_j}^{(\bar{j}+1)d_j} (x * h_h)[i, j, l] w_h[l] + n_h[\bar{i}, \bar{j}, l]$

$$y_h = \underbrace{S C_I W_h C_h T a}_{= H, \text{ forward spectrometer model matrix}} + n_h$$



Acknowledgments

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