



Instrumental development and preparation to future missions for Terrestrial Gamma ray Flash (TGF) detections



Mélody Pallu^{1,2} (pallu@apc.in2p3.fr), Philippe Laurent^{2,3}, Damien Pailot², Éric Bréelle², Sylvie Blin²

¹CNES, ²APC, ³CEA

1 – Overview

What are Terrestrial Gamma ray Flashes?

Bursts of gamma rays produced in thunderstorms ~400,000 TGFs/year detectable by satellites

~1 ph/cm² observed by satellite

TGF 081025

Duration: <100 µs

Light curve of a TGF detected by Fermi-GBM [Briggs et al., 2010].

- Photon energies: tens of keV to >40 MeV
- Very bright: 10¹⁸ photons produced

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- TGF current detections Mostly by satellites:
 - Mostly astrophysics instruments (e.g., Fermi, Agile, RHESSI)
 - Some TGF design instruments (ASIM on the ISS)
 - Some ground-based detections (with the Telescope Array)
- 2 aircraft detections

APC and CNES context

- Taranis was the 1st satellite designed for TGF study
- XGRE: X and gamma scintillator
- Launch failure in 2020

Objective: Develop an innovative gamma ray spectrometer multi-mission and for TGF detection Space instrument adaptable to detect different gamma ray events: e.g., TGFs, GRBs, solar bursts



3 – Future missions

* TGF X Nanosatellite

Short term: Balloon flight

- Planned in June 2024 in Kiruna, Sweden with CNES
- Stratospheric balloon staying at ~30 km for ~10 hrs
- To validate FGS working in conditions close to space: detection of Crab pulsar

Long term example: Bursty Energetic Events in Space (BEES)

- Nanosatellite constellation to study TGFs and GRBs
- Payload: FGS and a radio antenna from Czech Rep. team Aim: multiple TGF detection to study their characteristics, not assessable with only one measurement

Parallel study: Simulation of nanosatellite constellation for TGF detection

- TGF density map Nanosatellite trajectories Results of Monte-Carlo simulation of photon propagation in the atmosphere Objectives:
- - Determination of best nanosatellite configuration for TGF study Find a method to determine tilt angle, opening angle, and other TGF characteristics