



# Direct detection of exoplanets with MIRI on the James Webb Space Telescope

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<u>Challenge:</u> Direct imaging requires very high contrast at small angular separations.

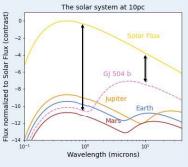


### Mid IR advantages:

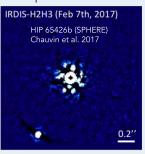
- Reduced contrast between the star and the planet
- Peak of emission of young giant planets
- Relevant for the detection of molecular signatures present in their atmospheres

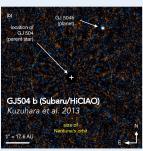
### Favorable cases

- → Young giant planets
- ightarrow Distant from their host star
- → IR Observations



<u>Solution:</u> Coronagraphic images (star hidden behind a mask) of young, long-period giant planets with an 8-meter telescope.





- → Only in near IR
- $\rightarrow$  Still impacted by the deformation due to the Earth atmosphere

Observations required from space in mid IR

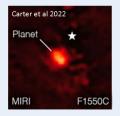
### <u>JWST</u>: 4 instruments including <u>MIRI</u> (Mid-IR Instrument)

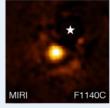
- $\rightarrow$  Spectral range from 5 to 28  $\mu$ m
- → First data available for direct imaging above 5 microns!

<u>Instruments</u>: Imager, Coronagraph, Low Resolution Spectrometer and <u>Medium Resolution Spectrometer</u>

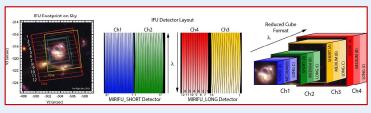
## 1 - MIRI coronagraphs:

First imaged exoplanet (HIP 65426 b)





#### 2 - MIRI Medium Resolution Spectrometer (MRS)

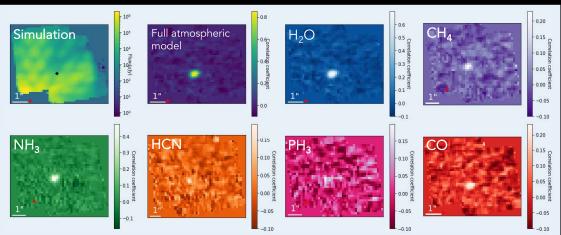


Integral field spectrometer: spectral and spatial information

ightarrow Allows us to distinguish spectrally and spatially the **star and the planet** because their spectra are different.

### "Molecular mapping" with MRS simulated data of the system GJ 504

- → Correlation with planetary atmospheric model spectrum and molecular template spectra
- $\rightarrow$  Increases the detection
- → Evidence for the presence of molecules



Diversity in the population of exoplanets: constraining the various planetary formation and evolution scenarios

- MIRI allows for the first detection of exoplanets in mid-IR.
- MRS will provide complementary data to characterize the atmosphere of young, giant planets.
- Detection of molecules that have previously never been confirmed in exoplanets (for example PH<sub>3</sub>) should be possible for cold planets at large separations (such as GJ 504 b).